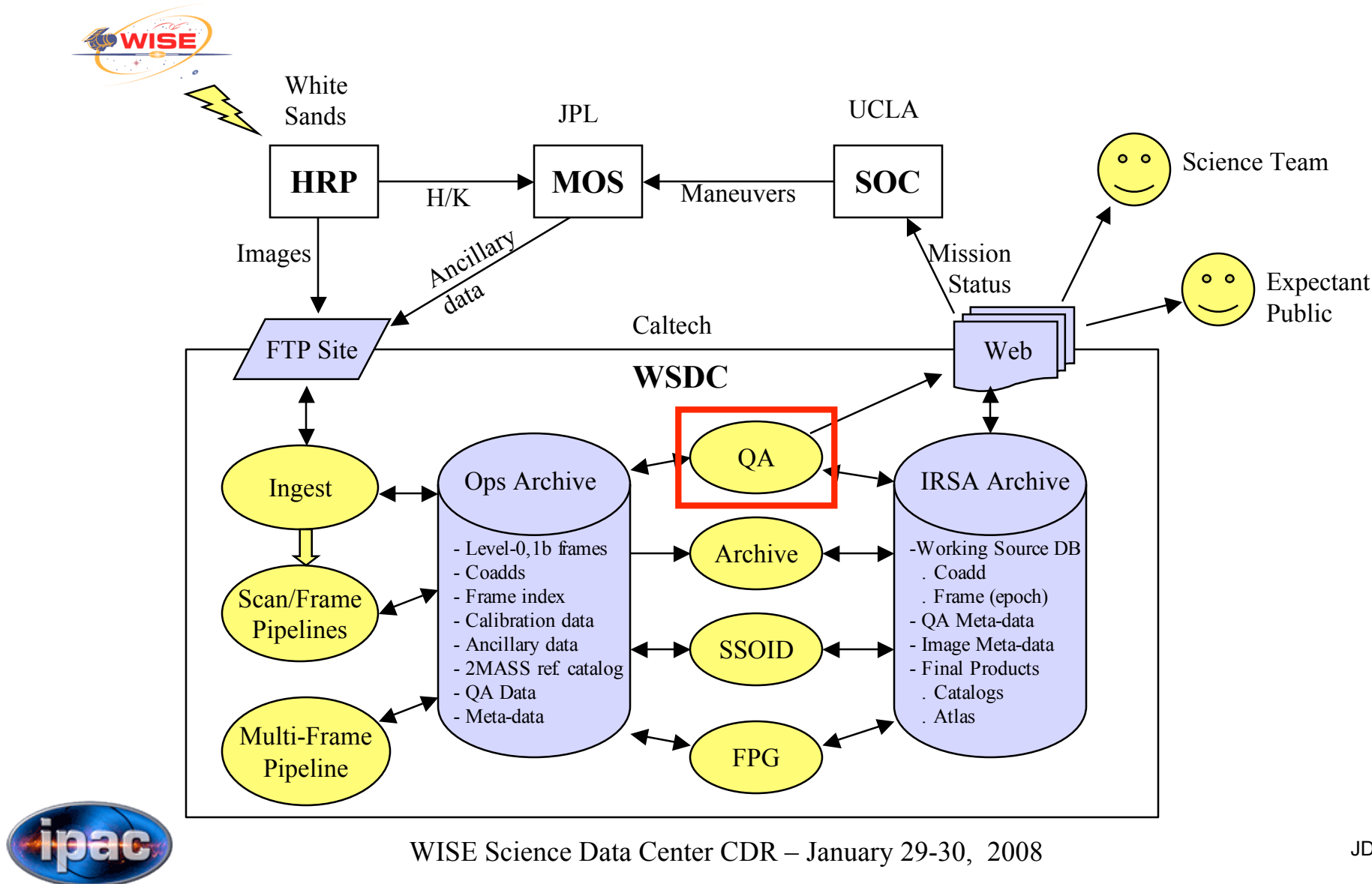


# Quality Assurance

Davy Kirkpatrick  
IPAC/Caltech



# WSDC Functional Block Diagram



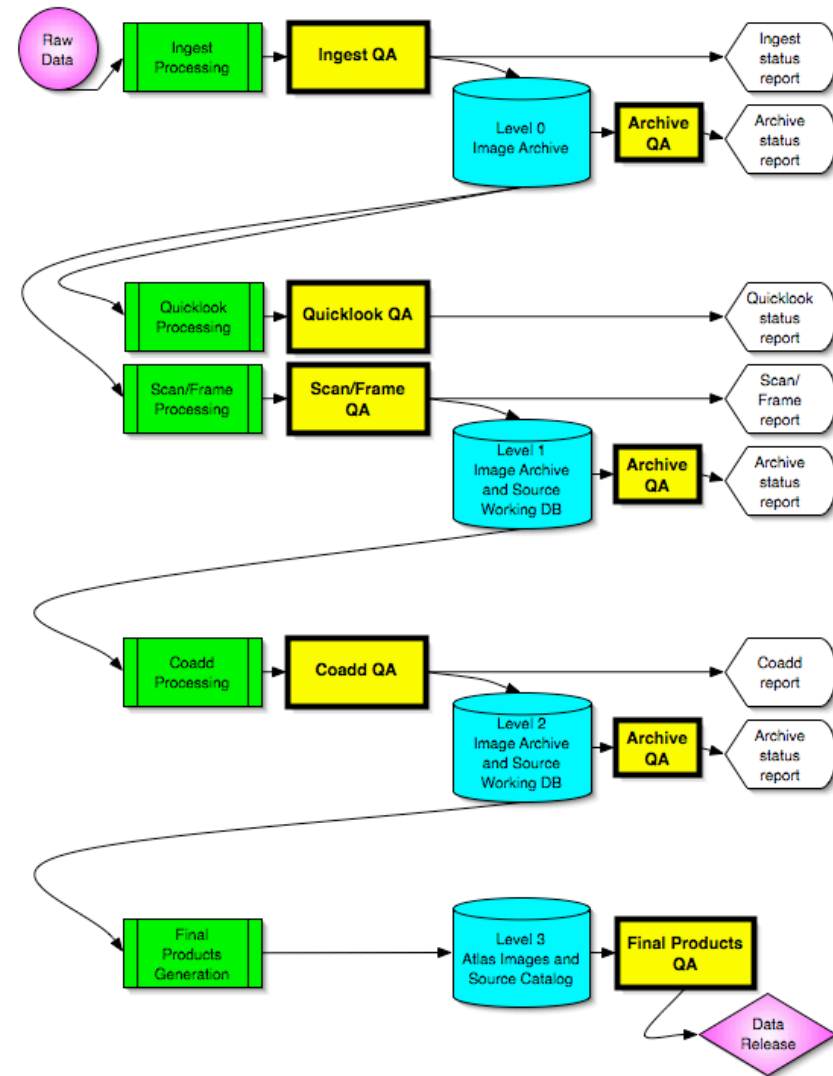


# QA Block Diagram



Quality Assurance

- Quality Assurance (QA) is injected into the data processing flow at many points to monitor data quality.
- QA subsystems are shown in yellow in the WSDC flow diagram at right. These are referred to as -
  - Ingest QA
  - Quicklook QA
  - Scan/Frame QA
  - Coadd QA
  - Archive QA
  - Final Products QA



# Driving Requirements

- L4WSDC-062: The WSDC shall perform quality analysis of all WISE science data and make reports available on a regular basis.
- L4WSDC-063: The WSDC shall work with the WISE Science Team to validate that, prior to their release, the Image Atlas and Source Catalog satisfy WISE science requirements.
- L4WSDC-064: The WSDC shall work with the WISE Science Team to characterize and document the overall data product relative to the mission requirements. This document shall be included in the WISE data product explanatory supplement.
- L4WSDC-065: A sample of 3% of the science imaging data returned to the ground each day shall be processed in an expedited way to produce a Quicklook report that monitors the routine performance of the flight system as can be determined from the science data, and identifies problems that may require prompt action by WISE Science or Mission Operations.
- L4WSDC-066: The WSDC shall provide a monitor of the synchronization between flight-system and scan mirror rates to achieve and maintain required image quality as part of Quicklook QA.

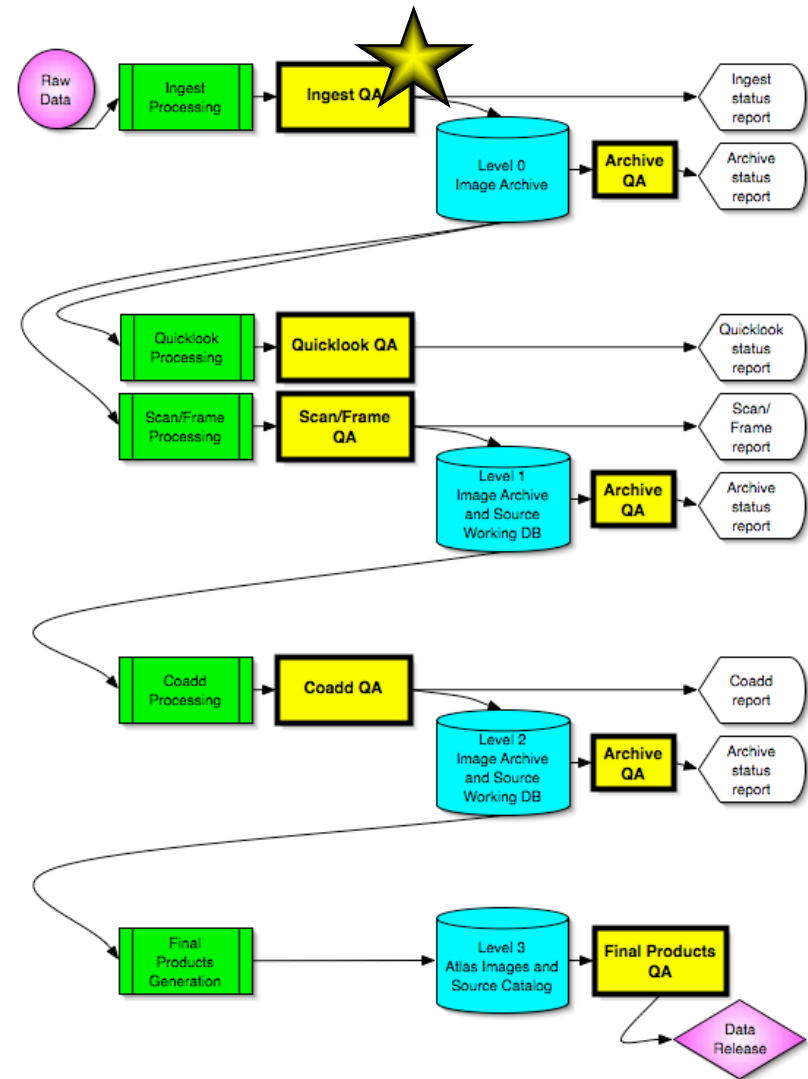


# Functionality: Ingest QA



Quality Assurance

- Purpose
  - To check compliance with the FITS standard.
  - To verify that all Level 0 images have been created.
  - To compare the input manifest from White Sands to the ingest output.
  - To stage data for further processing.
- Timescale:
  - Following each data transfer.
- Action:
  - WSDC to inform MOS/EOS and SOC of status and anomalies.



# Design: Ingest QA

- Check that assembled images meet the FITS standard.
- Compare input manifest to resulting output to check for completeness.
- Verify that all Level 0 images were successfully created.
- Verify that housekeeping telemetry data and SOE file were successfully mated with the correct images.
- Summarize QA findings for MOS and SOC.

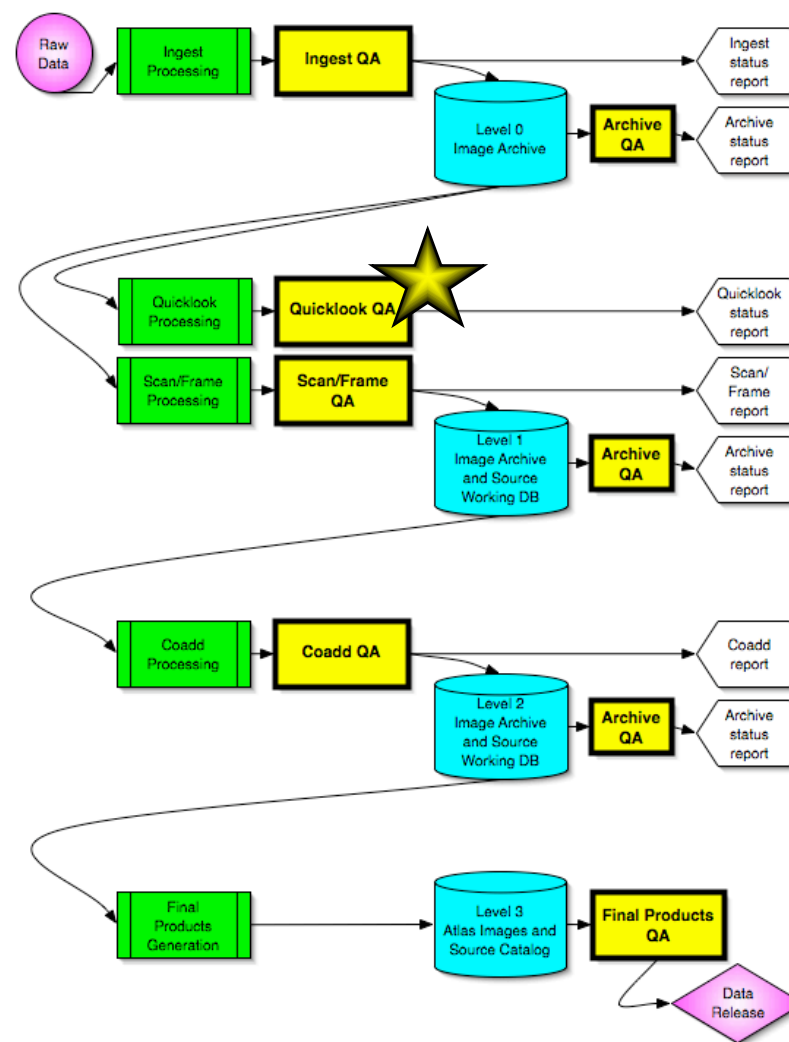


# Functionality: Quicklook QA



Quality Assurance

- Purpose:
  - To check system health (on 3% of data) for each downlink via an abbreviated processing pipeline. This will include monitoring of scan synchronization, system throughput, and image backgrounds and noise.
- Timescale:
  - Within 24 hours of end of data transfer to WSDC.
- Action:
  - WSDC to post report to web page; SOC to review report.





# Design: Quicklook QA (1)



- Scan synchronization and image quality
  - Generate matrix of composite star images detected on each frame in each band, where each element is the average star image formed by combining the images of all stars in the corresponding region of the frame.
  - Measure image second moment ratios and position angles for each composite star image in the matrix.
  - Generate table and plot showing the means of these values for all frames in a half-orbit.
  - Trigger warning messages when image elongation has exceeded a predetermined threshold related to the Level 1.5 specifications for image quality. Threshold values are to be determined prior to launch using simulated image data.
- Photometric zero point and system throughput (Needs polar data for primary and second standard star checks.)
  - Tabulate the mean and RMS differences between a priori “true” and measured instrumental magnitudes for standard stars observed in the polar frames.
  - Generate a table and plot showing the mean and RMS of these values for each orbit.
  - Trigger a warning message if the zero point offset in any band falls beyond a threshold value. The threshold value will be derived pre-launch and updated during IOC.







# Design: Quicklook QA (2)



- Image backgrounds and noise (requires polar data to use as bellwethers of background level)
  - Compute the mean pixel values along with total and point-source-filtered noise values for each frame.
  - Generate a table and plot of mean pixel values and noise levels for each frame in a half-orbit and for each quadrant (for Si:As arrays) or stripe (for HgCdTe arrays).
  - Compare the measured mean pixel values and noise values in each frame to threshold values, band by band.
  - Trigger a warning message if the mean pixel values and noise values exceed predefined thresholds. These thresholds will be determined pre-launch and updated during IOC.
- Visual checks
  - Generate jpegs of a few frames in each band and check by eye.
  - Generate three-color jpegs of a few registered framesets (if possible) and check by eye.

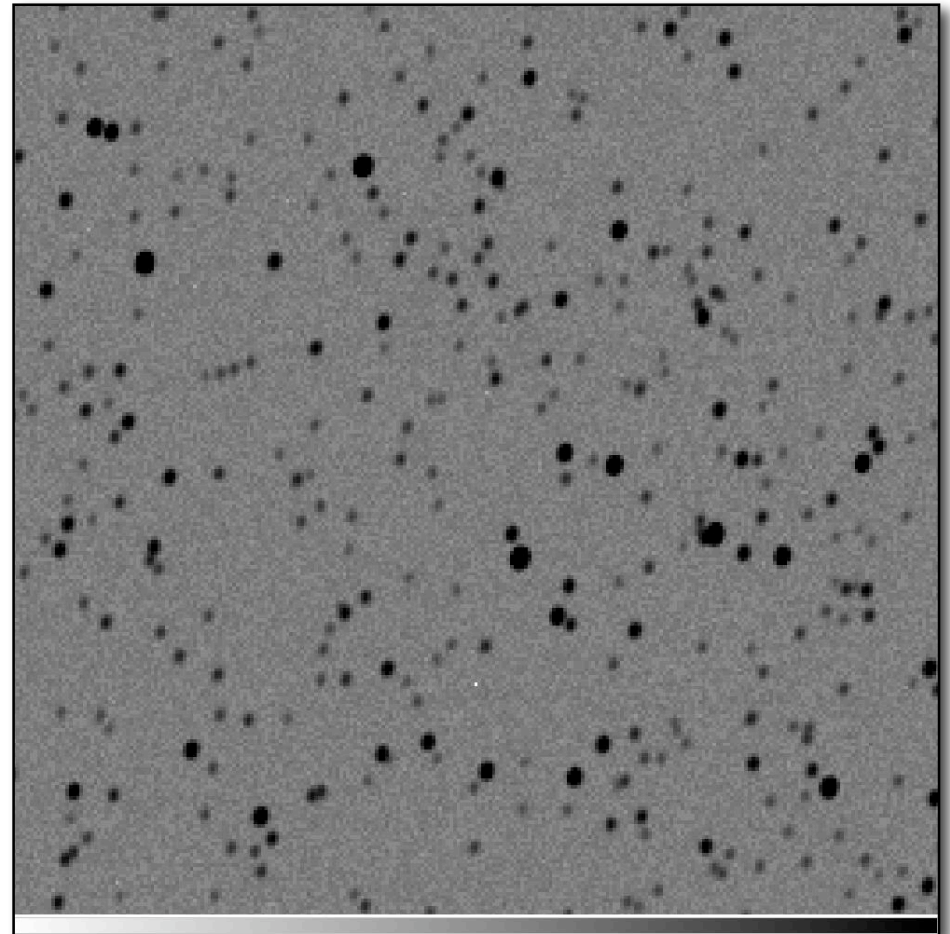




# Demo of Scan Synchronization Monitor (1)



- Image quality (via PRF) will be monitored as part of standard scan pipeline QA
- Same tool will be used to support IOC task to synchronize spacecraft/scan mirror rates
- Demonstrated using simulated IOC image data showing scan rate adjustments provided by Ned Wright
- Composite PRF generated by “stacking” high SNR point sources detected on one or more images
- PRF “shape” characterized, plotted as a function of frame, scan rate, etc.
- Optimal image shape occurs when scan rates are matched

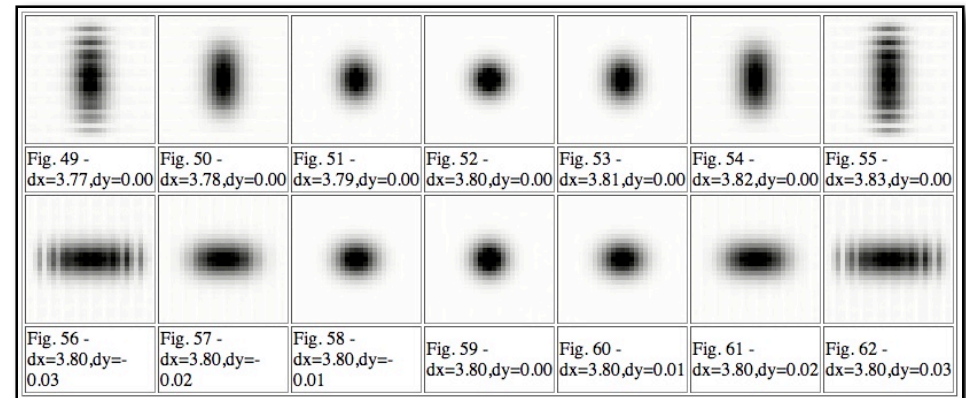




# Demo of Scan Synchronization Monitor (2)



- Image quality (via PRF) will be monitored as part of standard scan pipeline QA
- Same tool will be used to support IOC task to synchronize spacecraft/scan mirror rates
- Demonstrated using simulated IOC image data showing scan rate adjustments provided by Ned Wright
- Composite PRF generated by “stacking” high SNR point sources detected on one or more images
- PRF “shape” characterized, plotted as a function of frame, scan rate, etc.
- Optimal image shape occurs when scan rates are matched

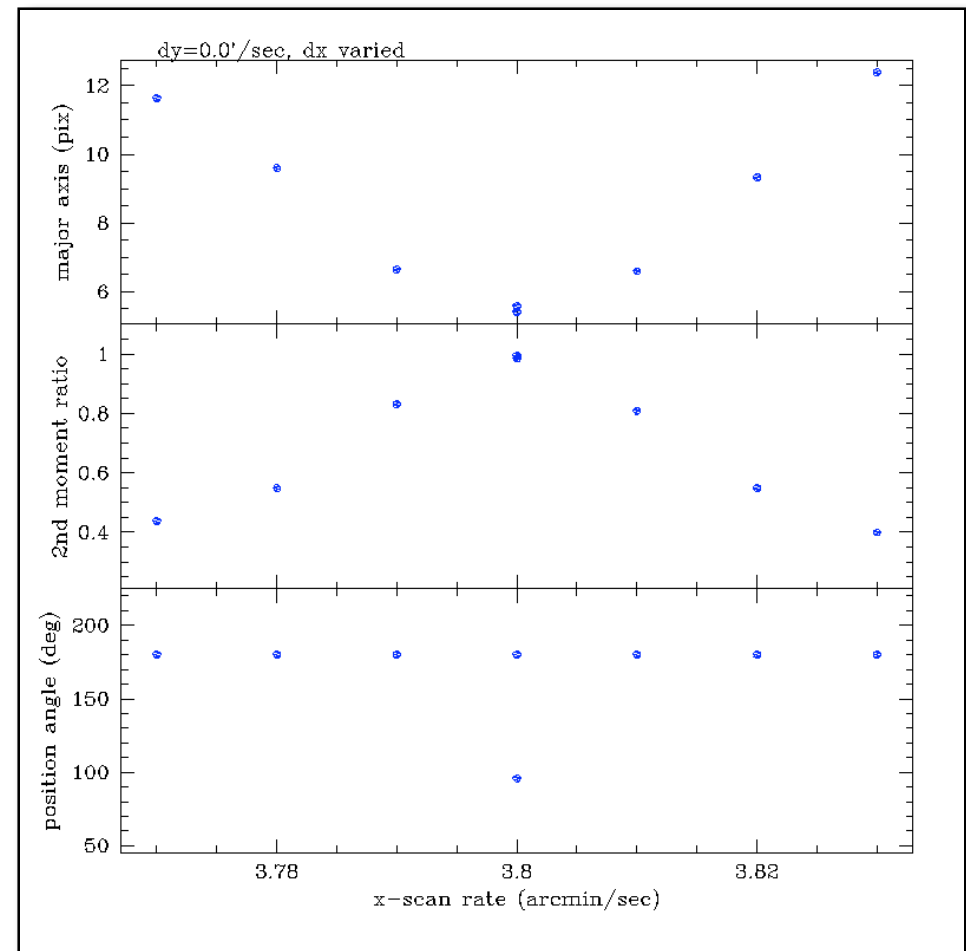




# Demo of Scan Synchronization Monitor (3)



- Image quality (via PRF) will be monitored as part of standard scan pipeline QA
- Same tool will be used to support IOC task to synchronize spacecraft/scan mirror rates
- Demonstrated using simulated IOC image data showing scan rate adjustments provided by Ned Wright
- Composite PRF generated by “stacking” high SNR point sources detected on one or more images
- PRF “shape” characterized, plotted as a function of frame, scan rate, etc.
- Optimal image shape occurs when scan rates are matched



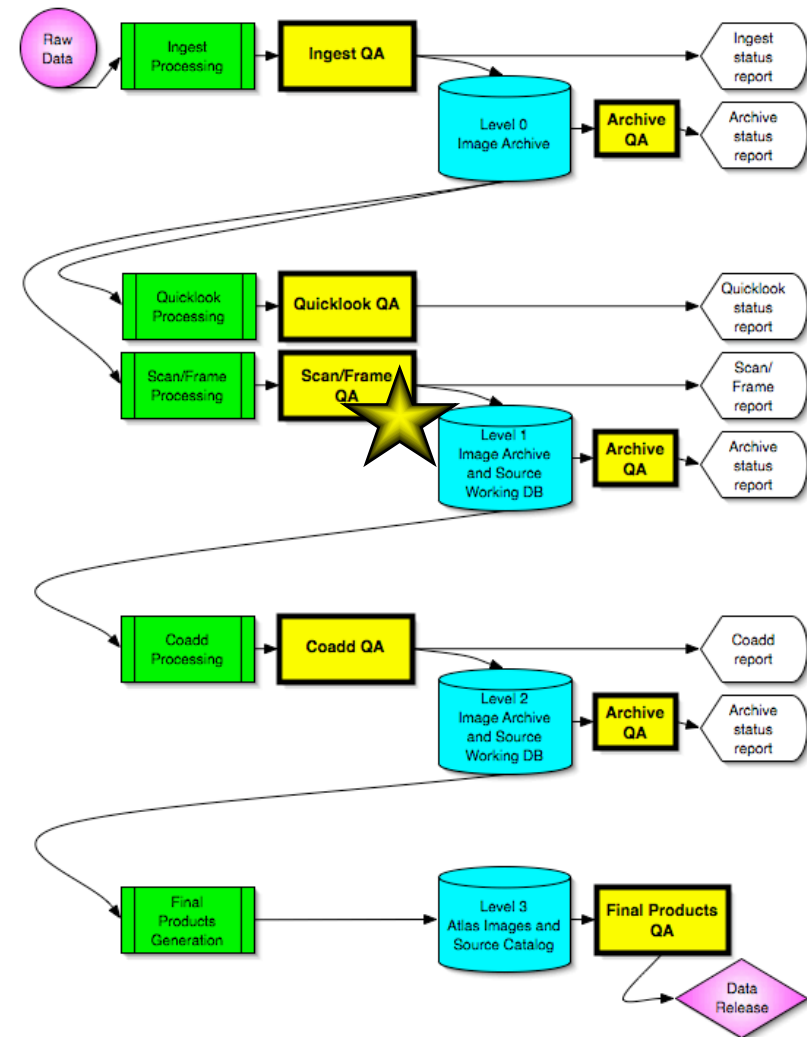


# Functionality: Scan/Frame QA



Quality Assurance

- Purpose:
  - To check for successful completion.
  - To scrutinize output of processing pipeline.
  - To compare performance to science metrics.
- Timescale:
  - Within 6 days.
- Action:
  - WSDC to assign quality scores to each frame or frameset and produce QA report; PI or his designee responsible for signing off.



# Design: Scan/Frame QA (1)

- Summary of input data
  - Report log file and results of ingestion QA.
  - Report QA results for quicklook processing.
- Instrumental image calibration
  - Compare flat-fields to fiducials.
  - Compare sky-offsets to fiducials.
  - Monitor dark images/overscans.
  - Monitor hot pixel masks – changes, # of pixels.
  - Monitor illumination profile corrections.
  - Flag outlying noisy frames; plot noise histograms.
  - Flag outlying point-source-filtered noisy frames; plot histograms.
- Scan synchronization
  - Monitor source shape, scan mirror synchronization.
- Bandmerging
  - Monitor band-to-band positional offsets.
  - Monitor percentage of sources seen in all bands vs. single-band missing sources, two-band missing, etc.



# Design: Scan/Frame QA (2)



- Astrometric calibration
  - Plot histograms of astrometric deltas between WISE-computed and 2MASS All-Sky PSC positions; scrutinize outliers.
  - Modulo solar system object identifications, tabulate and follow up sources with no 2MASS matches (at least in W1 and W2).
- Photometric calibration
  - Monitor mean aperture photometry curves-of-growth.
  - Tabulate/plot mean/RMS differences between truth and derived photometry for standard stars in the orbit.
  - Tabulate/plot mean/RMS differences between stars in this orbit and those observed in previous overlapping orbits (trending via other Level 1 data).
  - Tabulate/plot mean photometric offsets from in-scan overlaps.
  - Plot number of objects with noted source confusion as function of galactic latitude; spot check image data for selected clean and confused sources.
  - Plot saturated star mag/flux estimates against ramp saturation flag.
  - Compare color-color diagrams for objects saturated in any band and compare against fiducial color loci to check saturated mag estimates.







# Design: Scan/Frame QA (3)



- Artifact identification -- Perform semi-automated visual spot checks of a few examples of each of the following:
  - Latents.
  - Dichroic/filter glints.
  - Diffraction spikes.
  - Bright star halo contamination.
  - Optical ghosts.
  - Electronic ghosts.
  - Non-uniform stray light.
  - Scattered light patches from bright objects.
  - Radiation hits (?).
- Frame statistics
  - Plot  $\log(N)$ - $\log(S)$  and check against mean frame noise level.
  - Plot counts vs. ecliptic latitude as check for cosmic ray hits.
  - Measure frame-to-frame overlap to assure overlap is sufficient.







# Design: Scan/Frame QA (4)



- Solar system object identification
  - Plot number of solar system objects vs. ecliptic latitude.
  - Perform checks to make sure that identifications include asteroids, comets, planets, and planetary satellites.
  - Inspect color-color plots of identified objects.
  - Check detection fraction vs. visual magnitude?
- QA summary
  - Report successful/unsuccessful processing completion.
  - Provide web-accessible page with tables and plots listed above.
  - Generate auto-filled QA report along with quality scores as starting point for human review.
  - Review by QA scientists to finalize report.





# 2MASS Heritage: 2MASS Nightly QA (1)



2MASS Final Processing QA Review Status					
<div> <div>Online</div> <div>Ready for Review</div> <div>Being Reviewed</div> <div>Submitted</div> <div>Ready for Cleanup</div> <div>Pending Queries</div> <div>Northern Nights</div> <div>Southern Nights</div> <div>All Nights</div> </div>					
Night	Rev	Prod	Status	Reviewer	Review Status
980925a	QA	archived	010926	raymond	010924 approved 011001
980926a	QA	archived	011030	davy	011029 approved 011102
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980927a	QA	archived	010926	hurt	010926 approved 010928
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981012a	QA	archived	011102	hurt	011101 approved 011106
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981016a	QA	archived	011106	nelson	011104 approved 011106
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981102a	QA	archived	011106	hurt	011105 approved 011106

## 2MASS Pipeline 3 QA Report QA Web Index 981017s



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<a href="#">Photometricity</a>	<a href="#">Tracking/Stepping</a>	<a href="#">Observers Log</a>
<a href="#">Bkg/Seeing/Focus</a>	<a href="#">Jumps/Airglow</a>	<a href="#">V2 QA Web Pages</a>
<a href="#">Galaxies</a>	<a href="#">Nightly Overview</a>	<a href="#">V2 QA Review</a>

### Diagnostic Plots

<a href="#">Zero Point Default</a>	<a href="#">Seeing Shape &amp; Ratio</a>	<a href="#">Pos Reconst. Qual</a>	<a href="#">Chisq/Sig Plots</a>	<a href="#">Hess Plots</a>	<a href="#">Galaxy Monitor</a>
<a href="#">Overlaps Summary</a>	<a href="#">Meteor Blanking</a>	<a href="#">R1:R2 Diff Monitor</a>	<a href="#">PSF/Ap Mags</a>	<a href="#">In/Cross Scan Color</a>	<a href="#">V2 Deltas</a>
<a href="#">Sensitivity</a>	<a href="#">Astrom Wander</a>	<a href="#">Distortion Monitor</a>	<a href="#">R1:R2 Satur Phot</a>	<a href="#">Saturated Color-mags</a>	<a href="#">Plot Info</a>

### Error summary

filename	line #	error type	error message
input.c	2169	scan number not in list	data in updatesit for unknown scan#75 v2 grade=0
input.c	2169	scan number not in list	data in updatesit for unknown scan#76 v2 grade=0

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Version 010829 (3.1); more info is at [SCANSCLNotes](#)

Mail comments/bug reports to [hurt@ipac.caltech.edu](mailto:hurt@ipac.caltech.edu)





# 2MASS Heritage: 2MASS Nightly QA (2)

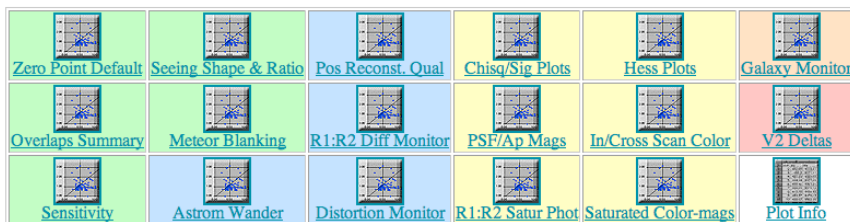


## 2MASS Pipeline 3 QA Report QA Web Index 981017s



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<a href="#">Photometricity</a>	<a href="#">Tracking/Stepping</a>	<a href="#">Observers Log</a>
<a href="#">Bkg/Seeing/Focus</a>	<a href="#">Jumps/Airglow</a>	<a href="#">V2 QA Web Pages</a>
<a href="#">Galaxies</a>	<a href="#">Nightly Overview</a>	<a href="#">V2 QA Review</a>

### Diagnostic Plots



### Error summary

filename	line #	error type	error message
input.c	2169	scan number not in list	data in updatesit for unknown scan#75 v2 grade=0
input.c	2169	scan number not in list	data in updatesit for unknown scan#76 v2 grade=0

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Last run: Wed Oct 31 15:38:25 2001

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Mail comments/bug reports to [hurt@ipac.caltech.edu](mailto:hurt@ipac.caltech.edu)

## 2MASS Pipeline 3 QA Report Quad Jump/Airglow/Planet Hazard 981017s



### Quadrant Jump/Banding Summary

Scan#	Type	Jump counter			Planet hrzds	Cnoise 3sig			Cnoise(4)			J Banding			H Banding			Ks Banding		
		J	H	Ks		J	H	Ks	J	H	Ks	L hi	R hi	L lo	L hi	R hi	L lo	L hi	R hi	L lo
007	CAL	0	0	0	0	undf	undf	undf	undf	undf	undf	0	0	0	0	0	0	0	0	0
008	CAL	0	0	0	0	undf	undf	undf	undf	undf	undf	0	0	0	2	0	0	1	0	0
009	CAL	0	0	0	0	undf	undf	undf	undf	undf	undf	0	0	0	0	0	0	0	0	0
010	CAL	0	0	0	0	undf	undf	undf	undf	undf	undf	0	0	0	1	0	0	1	0	0
011	CAL	0	0	0	0	undf	undf	undf	undf	undf	undf	0	0	0	0	0	0	0	0	0
012	CAL	0	0	0	0	undf	undf	undf	undf	undf	undf	0	0	0	1	0	1	0	2	0
013	SUR	0	0	0	0	0	0	0	0.500	0.200	0.400	0	0	0	2	0	0	1	2	0
014	SUR	0	0	0	0	0	0	0	0.600	0.200	0.400	0	0	1	2	0	0	2	1	0
015	SUR	0	0	0	0	0	0	0	0.800	0.300	0.300	1	0	3	2	0	0	2	1	0
016	SUR	0	0	0	0	0	0	0	0.600	0.300	0.500	0	0	1	2	0	0	2	0	1
017	SUR	0	0	0	0	0	0	0	0.600	0.100	0.500	0	0	1	1	0	0	0	3	0
018	SUR	0	0	0	0	0	0	0	0.500	0.200	0.400	0	0	0	2	0	0	2	1	0
019	SUR	0	0	0	0	0	0	0	0.600	0.100	0.300	0	0	0	5	0	3	3	10	0
020	SUR	0	0	0	0	0	0	0	0.600	0.200	0.300	0	0	0	2	0	0	2	3	0
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026	CAL	0	0	0	0	undf	undf	undf	undf	undf	undf	0	0	1	1	0	0	2	2	0
027	SUR	0	1	0	0	0	0	0	0.500	0.200	0.300	0	1	0	6	0	2	3	8	0
028	SUR	0	0	0	0	0	0	0	0.500	0.200	0.400	0	0	0	0	0	0	2	2	0





# 2MASS Heritage: 2MASS Nightly QA (3)



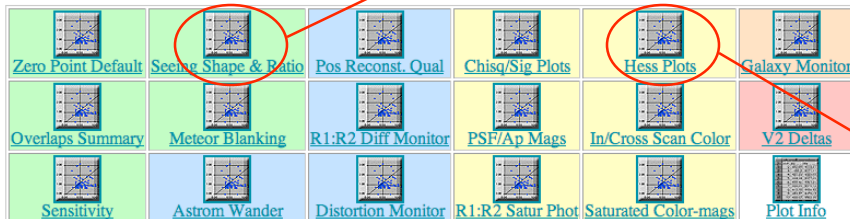
Quality Assurance

## 2MASS Pipeline 3 QA Report QA Web Index 981017s



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<a href="#">Photometricity</a>	<a href="#">Tracking/Stepping</a>	<a href="#">Observers Log</a>
<a href="#">Bkg/Seeing/Focus</a>	<a href="#">Jumps/Airglow</a>	<a href="#">V2 QA Web Pages</a>
<a href="#">Galaxies</a>	<a href="#">Nightly Overview</a>	<a href="#">V2 QA Review</a>

### Diagnostic Plots



### Error summary

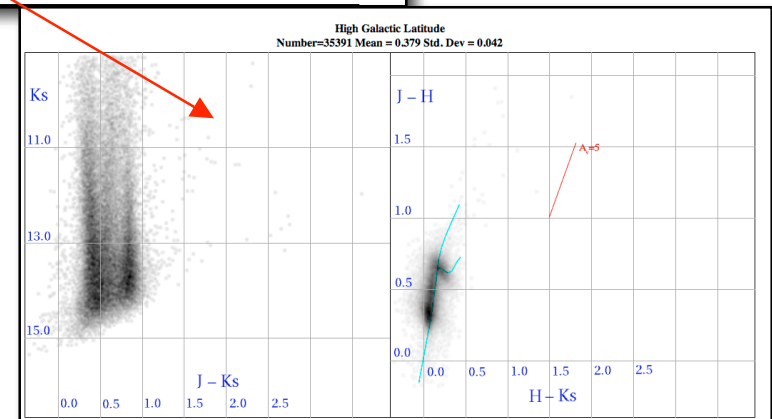
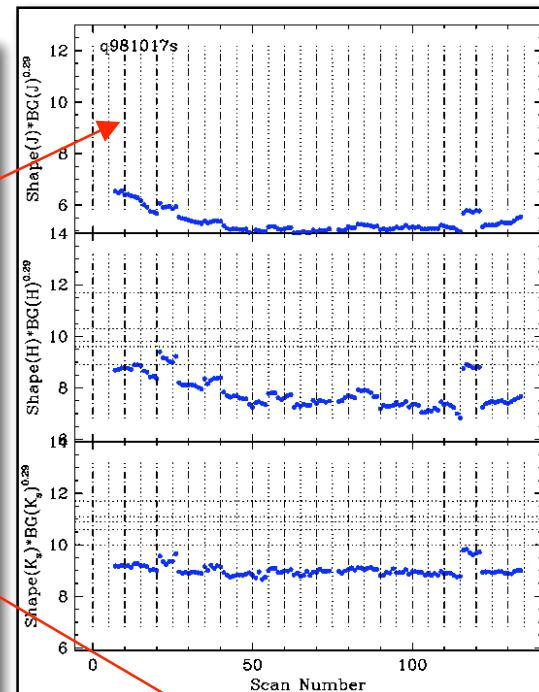
filename	line #	error type	error message
input.c	2169	scan number not in list	data in updatesit for unknown scan#75 v2 grade=0
input.c	2169	scan number not in list	data in updatesit for unknown scan#76 v2 grade=0

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# 2MASS Heritage: 2MASS Nightly QA (4)



Quality Assurance

<div> <div>Online</div> <div>Ready for Review</div> <div>Being Reviewed</div> <div>Submitted</div> <div>Ready for Cleanup</div> <div>Pending Queries</div> <div>Northern Nights</div> <div>Southern Nights</div> <div>All Nights</div> </div> <div>2MASS Final Processing QA Review Status</div>					
Night	Rev	Prod Status	Reviewer	Review Status	Other Notes
980925a	QA	archived 010926	raymond 010924	approved 011001	
980926a	QA	archived 011030	davy 011029	approved 011102	
980927a	QA	archived 020206	raymond 020206	approved	orig rev davy -----
980928a	QA	archived 010926	hurt 010926	approved 010928	position reconstruction problems i
980928a	QA	archived 020204	nelson 020204	approved 011001	orig rev raymond 010926
980929a	QA	archived 011031	nelson 011031	approved 011102	
980929a	QA	archived 010927	davy 010926	approved 011001	
980930a	QA	archived 011101	hurt 011101	approved 011105	
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981001a	QA	archived 010927	davy 010926	approved 011001	
981002a	QA	archived 011031	cxu 011029	approved 011101	
981002a	QA	archived 010927	davy 010926	approved 011001	
981003a	QA	archived 011031	cxu 011030	approved 011101	
981003a	QA	archived 010927	cxu 010926	approved 011002	
981004a	QA	archived 011030	cxu 011029	approved 011105	priority
981004a	QA	archived 010929	vandyk 010927	approved 011002	
981005a	QA	archived 011031	cxu 011030	approved 011105	
981005a	QA	archived 010927	vandyk 010926	approved 011002	
981006a	QA	archived 011031	davy 011030	approved 011105	priority
981006a	QA	archived 020206	nelson 020206	approved 011002	orig rev davy 010928
981007a	QA	archived 011031	nelson 011030	approved 011105	
981007a	QA	archived 011011	raymond 010927	approved 011015	
981008a	QA	archived 011031	cxu 011030	approved 011106	
981008a	QA	archived 020206	nelson 020206	approved 011002	orig rev raymond 010928
981009a	QA	archived 011031	cxu 011031	approved 011106	
981009a	QA	archived 020206	nelson 020206	approved 011002	orig rev davy 010928
981010a	QA	archived 011101	nelson 011031	approved 011106	
981010a	QA	archived 010928	cxu 010927	approved 010928	priority
981011a	QA	archived 011106	raymond 011102	approved 011106	
981011a	QA	archived 011002	hurt 011001	approved 011002	priority
981012a	QA	archived 011102	hurt 011101	approved 011106	
981012a	QA	archived 011002	hurt 011002	approved 011002	
981013a	QA	archived 011101	nelson 011071	approved 011106	
981014a	QA	archived 011102	hype 011101	approved 011107	
981014a	QA	archived 011001	hurt 010927	approved 011003	
981015a	QA	archived 011109	nelson 011104	approved 011107	rerun after dropping scan 098 from
981015a	QA	archived 020203	davy 011002	approved 011003	priority
981016a	QA	archived 011106	nelson 011104	approved 011106	
981016a	QA	archived 011002	vandyk 010928	approved 011004	
981017a	QA	archived 010929	davy 010928	approved 011004	
981018a	QA	archived 011003	cxu 011002	approved 011004	
981019a	QA	archived 011102	hurt 011102	approved 011107	priority
981019a	QA	archived 011002	raymond 010928	approved 011004	
981020a	QA	archived 011102	nelson 011102	approved 011107	
981020a	QA	archived 011001	cxu 011001	approved 011004	
981021a	QA	archived 011003	davy 011003	approved 011004	
981022a	QA	archived 011002	cxu 011001	approved 011004	priority
981022a	QA	archived 011109	hurt 011105	approved 011107	priority
981023a	QA	archived 011003	davy 011002	approved 011004	priority
981024a	QA	archived 011003	hurt 011001	approved 011002	priority
981025a	QA	archived 011002	cxu 011001	approved 011008	
981026a	QA	archived 011003	hurt 011002	approved 011008	
981027a	QA	archived 011106	nelson 011105	approved 011106	
981027a	QA	archived 011003	davy 011003	approved 011015	
981028a	QA	archived 020206	hurt 020204	approved	orig rev nelson -----
981028a	QA	archived 011017	hurt 011003	approved 011018	some star halo meteor blanking art
981029a	QA	archived 011103	hurt 011102	approved 011107	
981029a	QA	archived 011004	davy 011003	approved 011009	
981030a	QA	archived 020205	hurt 020204	approved 011009	
981031a	QA	archived 011001	cxu 011001	approved 011002	
981101a	QA	archived 011106	hurt 011105	approved 011106	
981101a	QA	archived 011003	raymond 011001	approved 011015	
981102a	QA	archived 011106	hurt 011105	approved 011106	

981017a [v3]  
This night has two scans that will need to be flagged to assure that their noted astrometric problems are factored into their final astrometric errors.

[http://2massqa.ipac.caltech.edu:9000/data/2mass/lgos\\_v3/qa981017a/](http://2massqa.ipac.caltech.edu:9000/data/2mass/lgos_v3/qa981017a/)

Processing notes:

Scans failing processing: None.

Minor planet ID's: Two low-numbered asteroids predicted and both found --

scan	psID	Name	RA	Dec
014	8521 (238)	Hypatia	329.664377	-6.891632
033	5658 (212)	Medea	331.222679	-8.601501

Photometricity:

Calibration strategy: New

CALMON overrides: None.

CALMON passes entire night with a solution in all three bands that approximates a constant fit.

Overlaps plots: Scatter is low in all sci blocks. Biases are near zero at J and H but are positive, though only at the 0.01 mag level, at Ks.

In summary,

Scans	Phot. quality factor
013-020	1.0
027-034	1.0
041-048	1.0
055-062	1.0
069-074	1.0
083-089	1.0
096-102	1.0
109-115	1.0
122-128	1.0

Backgrounds/Coads:

Background plots: These all look very stable. (I love southern data!)

Meteor blanking: Scan 015 (fr 126) contains a meteor with three parallel trails, and these were all cleaned successfully. A partial-frame meteor in scan 016 (fr 39) is also cleaned out completely. Scan 017 (fr 124) has two parallel meteors, only one of which is caught; the one missed is so faint that it does not create any false sources. Scan 018 (fr 144), scan 061 (fr 243), and scan 113 (fr 253) contain partial-frame meteors which are successfully cleaned.

Jump Counters: All OK

Airglow/Coad Noise: All OK

Marginal Airglow/Coad Noise: All OK

Seeing Diagnostics:

Extensive Untracked Seeing: All OK

Marginal Seetracker Scores: All OK

Bad 2nd image moment ratios: All OK

Astrometry Diagnostics:

Astrometric wander [blue]: The dec.dpos file for 014 vs 015 shows that RA is displaced by 0.3" for the entire run of Decs. Interestingly, the dec.dpos files for the two pairs on either side of this (013 vs 014 and 015 vs 016) show delta RA's which sit perfectly on zero. Howard has looked at comparisons with UCAC7, and it appears that both scans (014 and 015) are discrepant by about the same amount. Both of these scans will need to be flagged so that their final astrometric errors reflect this. This is by far the largest wander seen on this night; all others look fine.

Global astrometry [blue]: These look fine.

Scan tracking warnings: NONE

Science Diagnostics:

Photometric diagnostics [yellow]: Photometric sigmas show the usual families. Chi-squared values sit very close to 1.0 for all mags. Photometric linearity look very good except for the ~0.02 mag bias at the eastern edge of the array for faint Ks sources. All other plots look fine.

Galaxy diagnostics [orange]: These look fine.

Quality Recommendation:

Sci scans: (max score = 10; using min of fct's)

Scans	Grade	fct1	fct2	fct3	fct4	fct5
		phot	sens	see	untr	airy
013-020	10	1.0	1.0	1.0	1.0	1.0
027-034	10	1.0	1.0	1.0	1.0	1.0
041-048	10	1.0	1.0	1.0	1.0	1.0
055-062	10	1.0	1.0	1.0	1.0	1.0
069-074	10	1.0	1.0	1.0	1.0	1.0
083-089	10	1.0	1.0	1.0	1.0	1.0
096-102	10	1.0	1.0	1.0	1.0	1.0
109-115	10	1.0	1.0	1.0	1.0	1.0
122-128	10	1.0	1.0	1.0	1.0	1.0

Cal scans: (non-photometric = 0, photometric = 10)

Scans	Grade	Notes
007-012	10	
021-026	10	
035-040	10	
049-054	10	
063-068	10	
077-082	10	
090-095	10	
103-108	10	
116-121	10	
129-134	10	

Untracked Seeing Overrides: NONE

Missing Scans:

Scans	Notes
075-076	v2 Q=0

Version 2 Differences:

Delta Plots [red]: Zero-points show only expected changes. PSP values show almost no change at all. Source counts change by up to 60% but are as usual anti-correlated with source density.

Degraded scores: NONE

Improved scores: NONE

Special Notes:

None.

SCANSCI Version 010829 (3.1) last run on Fri Sep 28 06:33:13 2001

Final QA grades



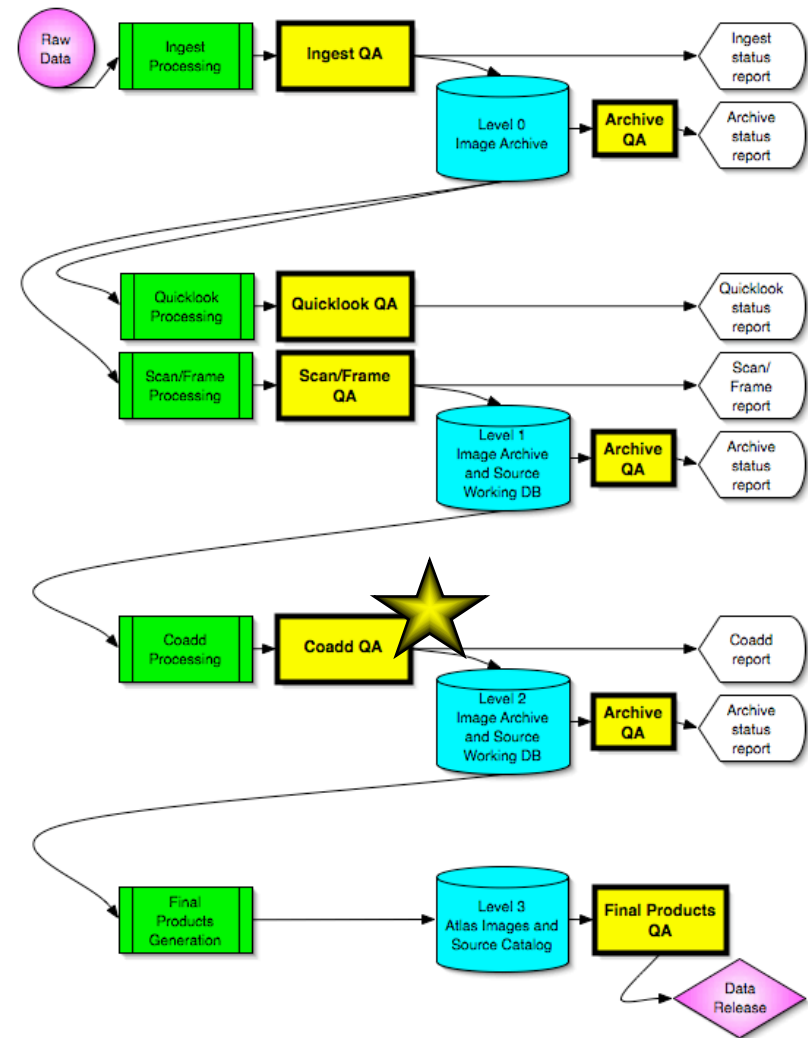


# Functionality: Coadd QA



Quality Assurance

- Purpose:
  - To check for successful completion.
  - To scrutinize output of processing pipeline.
  - To compare performance to science metrics.
- Timescale:
  - Within 15 days for multi-orbit runs with >18 coverages.
- Action:
  - WSDC to assign quality scores to each coadd or coaddset and produce QA report; PI or his designee responsible for signing off.



# Design: Coadd QA (1)

- Summary of input data
  - Summarize QA grades for each half-orbit considered for image stacking.
- Instrumental image calibration
  - Monitor changes and numbers of pixels in hot pixel masks?
  - Monitor illumination profile corrections?
  - Flag outlying noisy coadds; plot noise histograms.
  - Flag outlying point-source-filtered noisy coadds; plot histograms.
- Source characterization
  - Monitor source shape.
  - Perform semi-automated visual checks of registered coadds.
- Bandmerging
  - Monitor scan-to-scan positional offsets.
  - Monitor percentage of sources seen in all scans vs. single-scan missing sources, two-scan missing, etc.
- Astrometric calibration
  - Plot deltas with respect to 2MASS and individual scan (Level-1) astrometry; scrutinize sources with large deltas.
  - Modulo solar system object identifications, tabulate and follow up sources with no 2MASS matches (at least in W1 and W2).
  - Plot astrometric error per axis as function of source SNR

# Design: Coadd QA (2)

- Photometric calibration
  - Tabulate/plot zero-point differences for scan-to-scan overlaps.
  - Plot Level-1 photometry vs. deep coadd photometry to check for photometric self-consistency and depth of extractions.
  - Plot number of objects with noted source confusion as function of galactic latitude; spot check image data for selected clean and confused sources.
  - Plot saturated star mag/flux estimates against ramp saturation flag.
  - Compare color-color diagrams for objects saturated in any band and compare against fiducial color loci to check saturated mag estimates.
- Artifact identification
  - Check additional artifact flagging from extra-scan info; perform visual spot checks of a few examples of each artifact listed in the Scan QA list.
- Frame detection statistics
  - Plot  $\log(N)$ - $\log(S)$  and check against mean scan noise level or SNR.
  - Plot counts vs. ecliptic latitude as check for cosmic ray hits.
  - Measure scan-to-scan overlap to assure overlap is sufficient.



# Design: Coadd QA (3)

- Solar system objects identification
  - Plot number of solar system objects vs. ecliptic latitude.
  - Inspect color-color plots of identified objects.
  - Plot histogram of WISE magnitudes of newly identified sources (not found in Level 1 archive).
- QA summary
  - Report successful/unsuccessful processing completion.
  - Provide web-accessible page with tables and plots listed above.
  - Generate auto-filled QA report along with quality scores as starting point for human review.
  - Review by QA scientists to finalize report.

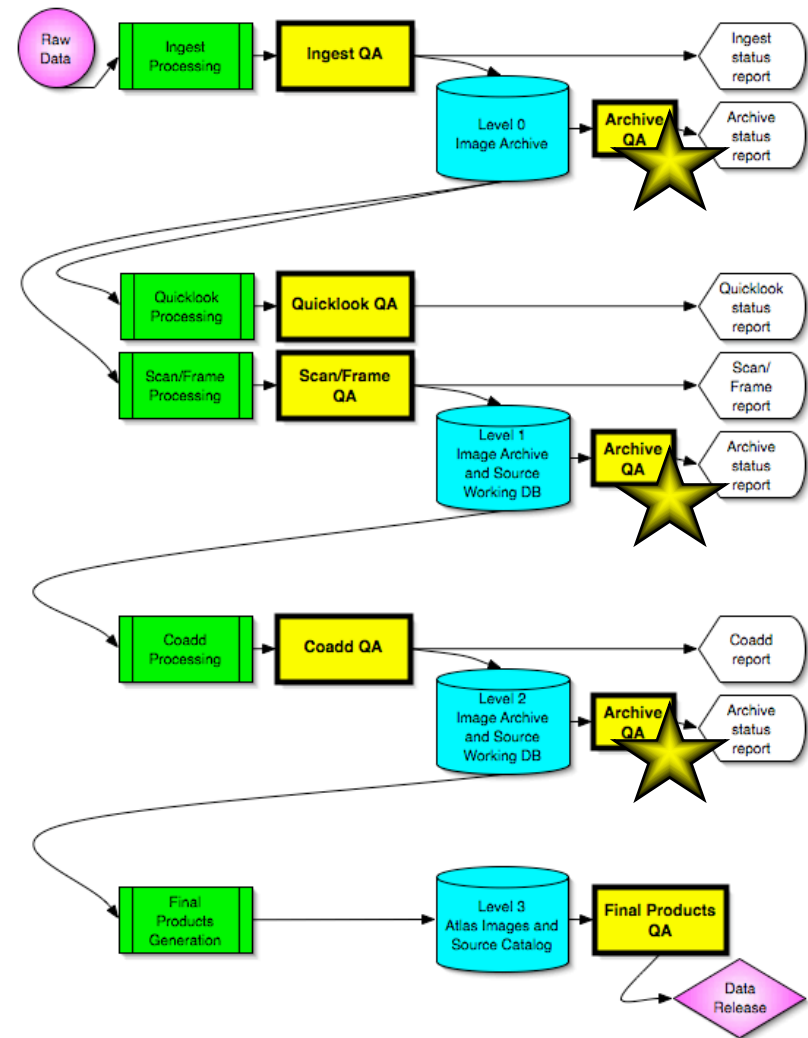


# Functionality: Archive QA



Quality Assurance

- Purpose:
  - To verify integrity of database tables and images (using, e.g., checksums and RTB queries).
  - To validate accuracy of source/metadata database loadings.
- Timescale:
  - After each database load.
  - Run periodically on static tables.
- Action:
  - WSDC reports status of checks and responds to problems (in concert with IRSA, where applicable).



# Design: Archive QA

- Perform checksums on the following:
  - Working databases.
  - Source catalogs.
  - Image archives.
  - Image metadata.
  - Any ancillary archives such as QA score archive or calibration archive.
- Perform range checking of the same databases, catalogs, and archives.
- QA summary
  - Report status of each check.
  - Provide web-accessible page with summarized results.

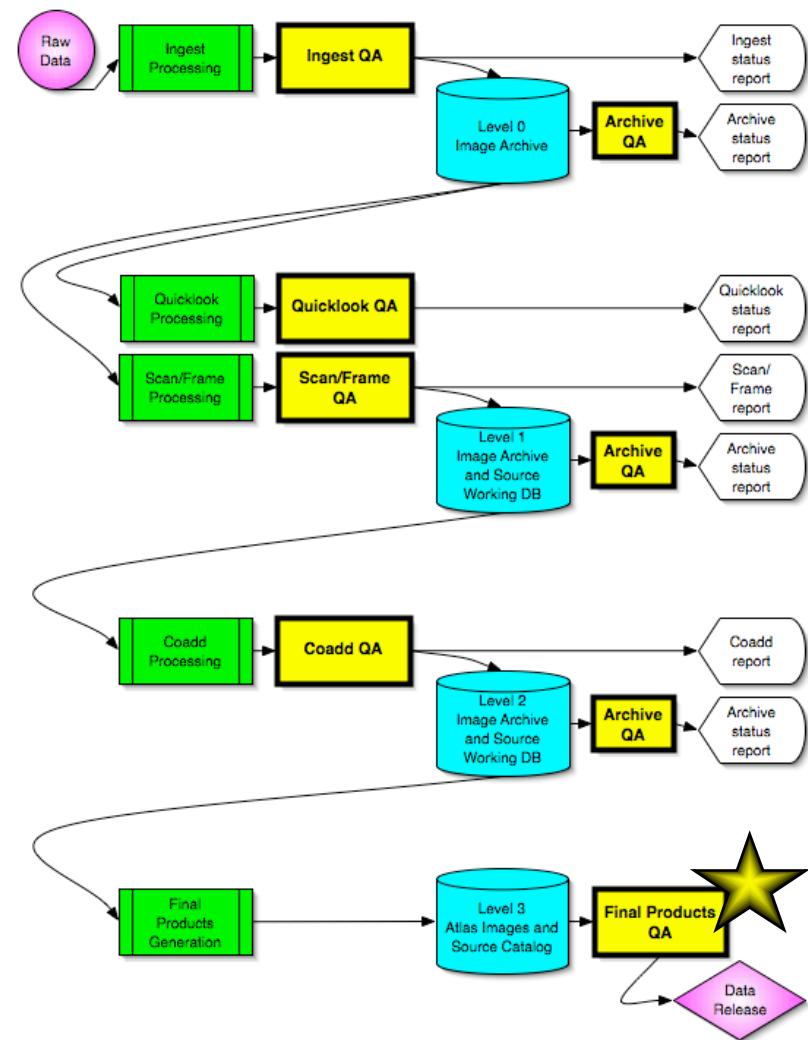


# Functionality: Final Products QA



Quality Assurance

- Purpose:
  - To assess properties of the Atlas Images and Source Catalogs relative to science requirements.
  - To check integrity of the products via range checking on all parameters.
  - To give overall characterization of public data products.
- Timescale:
  - After Final Product Generation but before public release.
- Action:
  - WSDC and Science Team to provide analyses; final release approval given by PI.



# Design: Final Products QA (1)

- Source Catalog
  - Plot histogram of N-out-of-M (N/M) statistics for multiple epochs; scrutinize cases with low N/M values.
  - Check photometric variability statistics; scrutinize outliers.
  - Check astrometric variability statistics (moving objects); scrutinize outliers.
  - Perform cross-correlations with other catalogs; scrutinize sources in each catalog that have no association in the other.
  - Perform range checking on all columns.
  - Using deeper Spitzer Space Telescope data at similar wavelengths (IRAC-ch1, IRAC-ch2, IRAC-ch4/IRS-blue-PU, and IRS-red-PU/MIPS-24um) and multi-repeat WISE scans at the ecliptic poles, determine completeness of catalog as a function of SNR and check with respect to science requirements.
  - Using the same data sets as above, determine reliability of catalog as a function of SNR and check with respect to science requirements.
  - Plot  $\log(N)$ - $\log(S)$  against mean scan noise level or SNR.
  - Plot saturated star mag/flux estimates against ramp saturation flag.
  - Compare color-color diagrams for objects saturated in any band and compare against fiducial color loci to check saturated mag estimates.
  - Plot astrometric error per axis as function of source SNR.



# Design: Final Products QA (2)



- Image Atlas (data images, depth-of-coverage mags, noise maps)
  - Confirm FITS standard.
  - Confirm that photometric zero points are correct.
  - Overlay images on outside image source data (2MASS All-Sky Atlas Images) to check astrometry.
  - Perform range checking on header values; check that pixel grid is the same for all images and all wavelengths.
  - Determine areal coverage for 8+-deep coverage areas and check against science requirement.
  - Using header info, determine epoch difference between first and last observation and check against science requirement.
  - Summarize reports of QA analyses for Explanatory Supplement.





# 2MASS Heritage: Final Products Analysis



	Category	Requirement	Performance
Point Source Catalog	Photometric sensitivity	10- $\sigma$ at 15.8, 15.1, 14.3 mag at J, H, K <sub>s</sub> respectively for $ \text{bl}  > 10^\circ$	met for full unconfused sky; exceeded for most of sky
	Photometric uniformity	<4% maximum bias in photometric zeropoint around the sky	<2% achieved
	Photometric precision	<5% 1- $\sigma$ for bright stars unsaturated in the 1.3 s exposure	<3% achieved
	Astrometric accuracy	<0.5'' 1- $\sigma$ relative to the reference frame	<0.1'' achieved
	Completeness	>0.99 at 10- $\sigma$ sensitivity limits	met
	Reliability	>0.9995	0.9997 demonstrated in test areas; no known source of unreliability in excess of specification
Bright star photometry (Read_1, rd_flg=1)		<2% bias at Read_1 saturation limit (K <sub>s</sub> ~4.0 mag)	<4% at worst
		<5% 1- $\sigma$ repeatability at Read_1 faint limit (K <sub>s</sub> ~8.0 mag)	2% achieved
		<10% 1- $\sigma$ repeatability at Read_1 saturation limit (K <sub>s</sub> ~4.0 mag)	2% achieved
Extended Source Catalog	Photometric sensitivity	10- $\sigma$ at J<15.0, H<14.2, K <sub>s</sub> <13.5 mag	met
	Photometric precision	<10% 1- $\sigma$ repeatability for H<13.8 mag	7-10% achieved
	Photometric uniformity	<10% maximum bias around sky	<4% achieved
	Completeness	>0.90 for $ \text{bl}  > 30^\circ$	met
	Reliability	>0.80 for $10^\circ <  \text{bl}  < 20^\circ$ >0.99 for $ \text{bl}  > 20^\circ$	achieved 0.93-0.95 achieved 0.992-0.995
General Survey	Sky coverage	>95%	99.998% imaged; 99.5% coverage in point sources; 98% coverage in extended source
Other Information	Galactic Plane performance	No Requirements	

## g. PSC Bright Source Photometric Precision

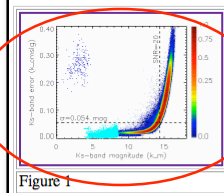
The Level 1 Specifications place three requirements on saturated star photometry (specifically "Read\_1" photometry).

1. 5% photometric precision for K<sub>s</sub>=8.0 mag (meaning just above the "Read\_2 - Read\_1" 1.3 s exposure saturation threshold (designated by rd\_flg=1)).
2. 10% photometric precision for at K<sub>s</sub>=4.0 mag (meaning just below the "Read\_1" 51 ms saturation threshold (thus, also rd\_flg=1)).
3. No more than 2% bias for K<sub>s</sub>>4.0 mag

There are no Level 1 requirements associated with sources that saturated even the 51 ms "Read\_1" exposure (designated by rd\_flg=3). These sources do have large uncertainties significantly in excess of the Level 1 Specifications for fainter sources.

## i. Bright Star Photometric Precision

Requirements 1 and 2 are addressed in the Photometric Precision subsection (see VI.1c), which discusses uncertainties derived from repeated observation across virtually the entire flux range observed in 2MASS. The relevant diagram is reproduced in Figure 1 and shows that the uncertainty in the magnitude range 4.0<K<sub>s</sub>[mag]<8.0 is substantially better than the 10%/5% requirements in items 1 and 2 above.

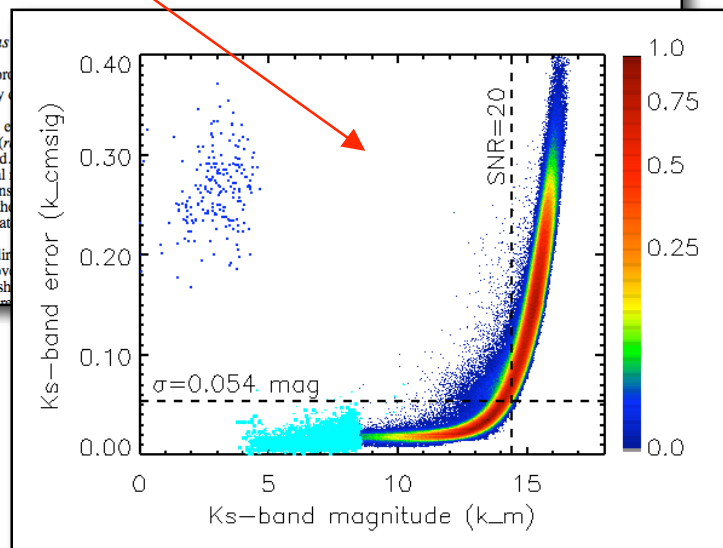


## ii. Bright Star Photometric Bias

A means has not been found to rigorously address the bias in the mag regime -- from internal Survey

Bias can be addressed at the "faint" end of the 51 ms Read\_1 photometry (rd\_flg=1) saturation threshold magnitude. Figure 2 shows a typical transition to saturated 1.3 s detections diagrams show deviation, because the deviation, because the singly-correlated

Although this result does not bear directly on the 1.3 s frames at the crossover, the linear right up to the saturation threshold, the linearity characteristics of this dev





# Anomaly Response Plan (1)



During the QA process, anomalies and problems will be detected that need response from the EOS, MOS, SOC, WSDC, or WISE Science Team.

- Ingest QA
  - Action: WSDC to inform MOS/EOS and SOC of anomalies via web-accessible report and by e-mail.
  - Tracking: MOS/EOS and/or SOC to acknowledge receipt of anomaly report via web form; results of subsequent analyses to be archived by WSDC and made web accessible.
  - Resolution: WSDC will close the issue and note it on web summary pages once parties agree that resolution has been reached.
- Quicklook QA
  - Action: WSDC to inform MOS/EOS and SOC of anomalies via web-accessible report and by e-mail.
  - Tracking: MOS/EOS and/or SOC to acknowledge receipt of anomaly report via web form; results of subsequent analyses to be archived by WSDC and made web accessible.
  - Resolution: WSDC will close the issue and note it on web summary pages once parties agree that resolution has been reached





# Anomaly Response Plan (2)

- Scan/Frame QA
  - Action: WSDC to report anomalies as part of the normal assigning of scan/frame QA grades. This report will be posted on the web for review by the PI and SOC.
  - Tracking: On the web summary pages, WSDC will mark reviews having outstanding issues and add subsequent analyses by WSDC or SOC to the QA review for curation.
  - Resolution: WSDC will close the issue and note it on web summary pages once the PI or designee concurs that resolution has been reached.
  
- Coadd QA
  - Action: WSDC to report anomalies as part of the normal assigning of coadd QA grades. This report will be posted on the web for review by the PI and SOC.
  - Tracking: On the web summary pages, WSDC will mark reviews having outstanding issues and add subsequent analyses by WSDC or SOC to the QA review for curation.
  - Resolution: WSDC will close the issue and note it on web summary pages once the PI or designee concurs that resolution has been reached.



# Anomaly Response Plan (3)



- Archive QA
  - Action: WSDC to report problems on web summary pages and to assign action items to internal WSDC or IRSA personnel via those project's ticket systems.
  - Tracking: Tracking will be handled by each project's existing ticket systems.
  - Resolution: WSDC will close issue and note it on web summary pages after concurrence of WSDC/IRSA personnel.
- Final Products QA
  - Action: WSDC to report anomalies to WSDC/IRSA via ticket systems or to SOC via e-mail.
  - Tracking: WSDC to note status of anomaly checking via ticket systems (for WSDC/IRSA-related issues) and via web summary pages.
  - Resolution: WSDC will close the issue and note it on web summary pages once the PI or designee concurs either that a solution has been found or that the anomaly should be characterized and documented for the Explanatory Supplement.





# QA Development Schedule



- Prototype QA - July 7, 2008
  - blah
- Preliminary QA - February 15, 2009
  - blah
- Complete QA - August 15, 2009
  - blah
- Mature QA - January 15, 2010
  - blah





# QA Staffing Schedule



Quality Assurance

- QA Lead Scientist
  - 0.3 (+0.2 as Sci Team member) FTE from now until ~6 months after final data release
- QA System Development Engineer
  - 0.5 FTE starting ~1.5yr before launch.
  - Ramps up to 1.0 FTE at ~6 months before launch.
  - Ramps back down to 0.5 FTE at ~2 months after end of on-orbit operations.
  - Ends at final data release.
- QA Scientists (numbers may include up to three postdocs)
  - 5.0 FTEs phased in between ~6 to ~9 months before launch and continue through final data release.
  - Ramps down to zero at ~6 months after final data release.
- Documentarian and User Support
  - 0.2 FTE starting ~3 months before launch.
  - Ramps up to 0.5 FTE at ~5 months after end of on-orbit ops (i.e., the start of final data processing).
  - Ends at ~6 months after final data release.



# Issues/Concerns

# Back-up Slides



# IOC: QA Tasks w/ Cover On



Although some of the data acquired during IOC will be non-standard with respect to normal operational data, QA tasks will be needed on resultant data products as the integrity of both the data and the pipelines are tested.

- Cover-on: Before the cover is released, the temperature of the inner shield of the aperture cover will be high enough that data in W3 and W4 should be saturated. Nonetheless, there are tests that can be made for W1 and W2 data:
  - Test Ingest QA pipeline.
  - Test those portions of the Scan/Frame QA Pipeline that monitor darks, hot pixel masks, and frame noise (for W1 and W2 only).
  - Check noise characteristics in W1 and W2 for orbits with SAA passages to determine noise thresholds to use during routine QA.





# IOC: QA Tasks w/ Cover Off



Cover-off: After the cover is ejected, a series of planned tests will evaluate the performance of the instrument to actual astronomical sources. These data will allow (a) threshold checking needed for nominal QA operations and (b) thorough testing of the scan synchronization monitor:

- Scan Synchronization monitor
  - Test fully.
  - Derive warning thresholds for QA.
- Detector calibration
  - Derive fiducial flats.
  - Derive fiducial masks (low-response pixels + hot pixels).
- Photometric calibration initialization
  - Monitor orbit-to-orbit pole passages to check zero-point stability of standard stars.
  - Monitor photometric stability using orbit-to-orbit overlaps.
- Source detection initialization
  - Determine optimal SNR thresholds for detections to meet completeness requirements.
- Bright source artifact mapping
  - Set thresholds for flagging latents, diffraction spikes, dichroic glints, electronic ghosts, and optical ghosts.
- Annealing characterization
  - Check image statistics before and after anneals.
  - Check behavior of latents before and after anneals.
- Avoidance limits
  - Determine practical background limit for processing data near the Moon.
  - Determine SAA charge rate limits for W3 and W4.

